Additional Problems For Self Practice (APSP)

PART -I : PRACTICE TEST PAPER

Max. Marks : 120

Important Instructions :

- 1. The test is of 1 hour duration and max. marks 120.
- 2. The test consists 30 questions, 4 marks each.
- 3. Only one choice is correct 1 mark will be deducted for incorrect response. No deduction from the total score will be made if no response is indicated for an item in the answer sheet.
- 4. There is only one correct response for each question. Filling up more than one response in any question will be treated as wrong response and marks for wrong response will be deducted accordingly as per instructions 3 above.
- 1. The density of newly discovered planet is twice that of earth. The acceleration due to gravity at the surface of the planet is equal to that at the surface of the earth. If the radius of the earth is R, the radius of the planet would be :

 $(3) \frac{1}{4}R$

- (1) 2R (2) 4R
- 2. Imagine a new planet having the same density as that of earth but it's radius is 3 times bigger than the earth. If the acceleration due to gravity on the surface of earth is g and that on the surface of the new planet is g', then :

(1)
$$g' = 3g$$
 (2) $g' = \frac{g}{9}$ (3) $g' = 9g$ (4) $g' = 27g$

3. Change in acceleration due to gravity is same upto a height h from earth surface and below depth x then relation between x and h is (h and $x \ll R_e$):

h

- 5. A geostationary satellite is rotating in circular orbit of radius 36000 km around the earth. A spy satellite which is rotating in circular orbit at a height of some hundred kilometre from earth's surface has time period approximately equal to ($R_e = 6400$ km): (1) 1 hour (2) 2 hour (3) 24 hour (4) 36 hour
- 6. Assuming earth to be sphere of uniform density what is the value of acceleration due to gravity at a point 100 km below the earth surface (Given R = 6380×10^3 m): (1) 3.10 m/s (2) 5.06 m/s² (3) 7.64 m/s² (4) 9.66 m/s²
- 7. The weight of an object at earth's surface is 700 g wt. What will be its weight at the surface of a planet whose radius is 1/2 and mass is 1/7 of that of the earth ?
 (1) 200 g wt.
 (2) 400 g wt.
 (3) 50 g wt.
 (4) 300 g wt.
- 8. A satellite moves around the earth in a circular orbit of radius r with speed v. If the mass of the satellite is M, its total energy is: (1) -1/2 Mv² (2) 1/2 Mv² (3) 3/2 Mv² (4) Mv²
- **9.** Three identical stars of mass M are located at the vertices of an equilateral triangle with side L. The speed at which they will move if they all revolve under the influence of one another's gravitational force in a circular orbit circumscribing the triangle while still preserving the equilateral triangle :

Max. Time : 1 Hr.

 $(4) \frac{1}{2}R$

Gravi	tation							
	2GM	GM	2 GM					
	(1) ^V L	(2) ^V ∟	(3) [−] V L	(4) not possible at all				
10.	In a spherical region, the a distance r from the co	ne density varies inverse entre is : 1	ly with the distance from	the centre. Gravitational field at				
	(1) proportional to r	(2) proportional to r	(3) proportional to r ²	(4) same everywhere				
11.	In above problem, the	gravitational potential is ·	- 1					
	 (1) linearly dependent (3) proportional to r² 	on r	 (2) proportional to r (4) same every where. 					
12.	A body of mass m is lift change in potential energy	red up from the surface overgy of the body is $(g = g)^2$	f earth to a height three t ravity field at the surface	imes the radius of the earth. The of the earth)				
	(1) mgR	(2) $\frac{3}{4}$ mgR	(3) ¹ / ₃ mgR	(4) $\frac{2}{3}$ mgR				
13.	A satellite can be in a g the angular velocity of around the planet if its	geostationary orbit aroun the planet about its axis distance from the centre	d a planet at a distance doubles, a satellite can of the planet is	r from the centre of the planet. If now be in a geostationary orbit				
	(1) $\frac{r}{2}$	(2) $\frac{1}{2\sqrt{2}}$	(3) $\overline{(4)^{1/3}}$	(4) $\overline{(2)^{1/3}}$				
14	Assuming the earth to gravitation) and g (accord) (1) 3g/(4πRG)	be a homogeneous sp eleration due to gravity o (2) 4πg/(3RG)	here of radius R, its de n the surface of the earth (3) 4πRg/(3G)	nsity in terms of G (constant of η) is (4) 4πRG/(3g)				
15.	Weight of an object is : (1) Normal reaction be (2) Gravitational force (3) dependent on frame (4) net force on the obj	tween ground and the ob exerted by earth on the c e of reference. ect	oject object.					
16.	In a binary star system of mass then : (1) Both stars have sar (2) Both stars have and (3) Both stars have sar (4) Both stars have sar	one star has thrice the m ne angular momentum a gular momentum of same ne angular speeds. ne linear speeds.	hass of other. The stars re bout common centre of r e magnitude about comm	otate about their common centre mass. non centre of mass.				
17.	An object is placed at a acceleration of that obj	a distance of R/2 from th ect due to gravity at that idius of earth)	e centre of earth. Knowi t point is: (g = accelerati	ng mass is distributed uniformly, ion due to gravity on the surface				
	(1) g	(2) 2 g	(3) g/2	(4) none of these				
18.	Altitude at which accele (Radius of earth = 640	eration due to gravity dec 0 km)	creases by 0.1% approxi	mately :				
	(1) 3.2 km	(2) 6.4 km	(3) 2.4 km	(4) 1.6 km				
19.	A particle of mass M is the gravitational potent 2GM	s situated at the centre of ial at a point situated at a 3GM	f spherical shell of mass a/2 distance from the cer 4GM	and radius a. The magnitude of htre, will be : GM				

 $\begin{array}{ccc}
\frac{2GM}{a} & \frac{3GM}{a} & \frac{4GM}{a} & \frac{GM}{a}
\end{array}$

20.	A satellite of mass m is circulating around the earth with constant angular velocity. If radius of the orbit is R_0 and mass of the earth M, the angular momentum about the centre of the earth is								
	(1) $m\sqrt{GMR_0}$	(2) $M\sqrt{GmR_0}$	(3) $m\sqrt{\frac{GM}{R_0}}$	(4) M	$\frac{GM}{R_0}$				
21.	Figure show a her of gravitational fie (1) a (3) c	nispherical shell havi ld intensity at point P	ng uniform mass dens will be along: (2) b (4) d	sity. The direction					
22.	Mass M is uniform hemispherical she of hemisphere, s potential at points (1) V _A > V _B >V _C	mly distributed only o II. <i>A</i> , <i>B</i> and <i>C</i> are thre uch that <i>A</i> is the c A, B and C be V _A , V _E Made MPS 20	on curved surface of e points on the circula entre. Let the gravit a, V _C respectively. The 08 (2) V _C > V _B >	a thin ar base rational en	A • B • C •				
	(3) $V_B > V_A$ and $V_B >$	Vc	(4) $V_A = V_B =$	Vc					
23.	A satellite close to above a point P on (1) 1.5 hours (2) 1.6 hours if it is (3) 24/17 hours if it (4) 24/17 hours if it	the earth is in orbit a the equator at some rotating from east to is rotating from east is rotating from west	bove the equator with time, it will be above l west to west to east	n a period of revolut P again after time	tion of 1.5 hours. If it is				
24.	The escape velocity (1) $\sqrt{\frac{2GM}{R}}$	y of a sphere of mass (2) $2\sqrt{\frac{GM}{R}}$	s m from earth having (3) $\sqrt{\frac{2GMm}{R}}$	mass M and radius $\sqrt{\frac{G}{F}}$	R is given by - M R				
25.	An artificial satellite	moving in a circular	orbit around the earth	has a total (kinetic	+ potential) energy E0.				
	Its potential energy	is :							
	(1) - E ^o (2)	1.5 E ^o	(3) 2 E ^o	(4) E ^o					
26.	A satellite S is mo compared to the ma (1) The acceleration (2) The angular mo remains constant (3) The total mecha	ving in an elliptical c ass of the earth : n of S is always direc omentum of S about unical energy of S var	orbit around the earth ted towards the centr the centre of the earth ies periodically with ti	h. The mass of the re of the earth th changes in direc me	e satellite is very smal tion, but its magnitude				

- (4) The linear momentum of S remains constant in magnitude.
- 27. If a body describes a circular motion under inverse square field, the time taken to complete one revolution T is related to the radius of the circular orbit as -(1) T \propto r (2) T \propto r2 (3) T2 \propto r3 (4) T \propto r4

Gravitation

- 28. Two satellites A and B go round a planet P in circular orbits having radill 4R and R respectivley. If the speed of the satellite A is 3V, the speed of the satellite B will be (1) 12 V (2) 6 V (3) 4/3 V (4) 3/2 V
- 29. A body attains a height equal to the radius of the earth. The velocity of the body with which it was projected is :

GM	2GM	5 GM	3GM
(1) √ R	(2) √ R	(3) V 4 R	(4) √ R

The distance of neptune and saturn from sun are nearly 1013 and 1012 meters respectively. Assuming 30. that they move in circular orbits, their periodic times will be in the ratio -(3) 10√10 (4) 1/√10 (2) 100

0110	1	2	3	4	5	6	7	8	٩	10
Que.	'		<u> </u>	-	,	v	,	Ů	J	10
Ans.										
Que.	11	12	13	14	15	16	17	18	19	20
Ans.										
Que.	21	22	23	24	25	26	27	28	29	30
Ans.										

Practice Test (JEE-Main Pattern)

PART - II : PRACTICE QUESTIONS

- The escape velocity from the earth is about 11 km/second. The escape velocity from a planet having 1. twice the radius and the same mean density as the earth is -(1) 22km/sec (2) 11 km/sec (3) 5.5 km/sec (4) 15.5 km/sec
- 2. A satellite which is geostationary in a particular orbit is taken to another orbit. Its distance from the centre of earth in new orbit is 2 times that of the earlier orbit. The time period in the second orbit is -

(2) $48\sqrt{2}$ hours (1) 4.8 hours (3) 24 hrs (4) Infinite

- A satellie revolves around the earth in an elliptical orbit. Its speed 3.
 - (1) Is the same at all point in the orbit
 - (2) Is greatest when it is closest to the earth
 - (3) Is greatest when it is farthest from the earth
 - (4) Goes on increasing or decreasing continuously depending upon the mass of the satellite
- The moon's radius is 1/4 that of the earth and its mass is 1/80 time that of the earth. If g represents the 4. acceleration due to gravity on the surface of the earth, that on the surface of the moon is (1) g/4 (2) g/5 (3) q/6 (4) q/8
- 5. A simple pendulum has a time period T_1 when on the earth's surface, and T_2 when taken to a height R above the earth's surface, where R is the radius of the earth. The value of T_2/T_1 is :

(1) 1 (2)
$$\sqrt{2}$$
 (3) 4 (4) 2

- The orbital speed of a satellite revolving nearby the earth is : 6. (4) $\sqrt{2g/R}$ (3) $\sqrt{g/R}$ (1) $\sqrt{2gR}$ (2) √gR
- 7. If the radius of earth is decreased by 1% and mass remain constant, then the acceleration due to gravity

Gravi	tation									
	(1) decrease by 2%	(2) decrease by 1%	(3) increase by 1%	(4) increase by 2%						
8.	The escape velocity of (1) m ²	a particle of mass m var (2) m	ies as : (3) mº	(4) m ⁻¹						
9.	Escape velocity for a ro to gravity is double than (1) 5.6 m/s	ocket is 11.2 km/s. If it is n earth, then escape velo (2) 11.2 m/s	taken to a planet where ocity will be : (3) 22.4 km/s	the radius and acceleration due						
10	If the change in the vel	(2) 11.2 1173	(0) 22.4 km/3	arth is the same as at a depth y						
10.	below it, then (both x and h being much smaller than the radius of the earth) - $x = \frac{h}{-}$									
	(1) x = h	(2) x = 2h	(3) * - 2	(4) x = h2						
11.	The period of a satellite of radius 4R is -	e in a circular orbit of radi	ius R is T, the period of a	nother satellite in a circular orbit						
	(1) 41	(2) 1/4	(3) 81	(4) 1/8						
12.	A spherical planet far acceleration due to gra	out in space has a ma vity which is equal to	ass M_0 and diameter D_0	b. A particle will experience an						
	(1) GM_0/D_02	(2) $2mGM_0/D_02$	(3) 4GM ₀ /D ₀ 2	(4) GMM_0/D_02						
13.	A geostationary satellit period of a spy satellite will approximately be :	e orbits around the eart orbiting a few hundred	h in a circular orbit of ra kilometers above the ea	adius 36000 km. Then, the time arth's surface (R _{Earth} = 6400 km)						
	(1) 1/2 hr	(2) 1 hr	(3) 2 hr	(4) 4 hr						
14.	The dimensions of univ (1) [M ⁻¹ L ³ T ⁻²]	versal gravitational consta (2) [ML ² T ⁻¹]	ant are : (3) [M ^{_2} L ³ T ^{_2}]	(4) [M ⁻² L ² T ⁻¹]						
15.	Suppose radius of the moon's orbit around the earth is doubled. Its period around the earth will bec									
	(1) 1/2 times	(2) $\sqrt{2}$ times	(3) 2 ^{2/3} times	(4) 2 ^{3/2} times						
16.	Acceleration due to gra 32 km from sea level : (1) 0.5 g ms ^{-2}	avity at earth's surface is ($R_e = 6400 \text{ km}$) (2) 0.99 g ms ⁻² (3) 1.0	s g ms ⁻² . Find the effectiv 1 g ms ⁻²	ve value of gravity at a height of (4) 0.90 g ms ⁻²						
17.	Near earth's time perio	d of a satellite is 4 h. Fin	d its time period at heigh	t 4R from the centre of earth :						
	·									
	(1) 32 h	(2) $\left(\frac{8^{3}\sqrt{2}}{9^{3}\sqrt{2}}\right)$ h	(3) ^{8 ³√2 h}	(4) 16 h						
18.	Radius of orbit of satell	ite of earth is R. Its kinet	ic energy is proportional	to :						
	(1) $\frac{1}{R}$	(2) $\frac{1}{\sqrt{R}}$	(3) R	(4) $\frac{1}{R^{3/2}}$						
19.	The radius of the orbit of (1) 4.2 T	of a planet is two times tl (2) 2.8 T	hat of the earth. The time (3) 5.6 T	e period of planet is : (4) 8.4 T						
20.	Two satellite of earth, S_2 . Which one of the form (1) The time period of S_2 . The potential energies (3) S_1 and S_2 are moving (4) The kinetic energies	S_1 and S_2 are moving in f_2 blowing statements is trues S_1 is four times that of S_2 ies of earth and satellite ing with the same speed s of the two satellites are	the same orbit. The mas ue : in the two cases are equ	s of S_1 is four times the mass of al						
21.	(4) The kinetic energies of the two satellites are equal A cosmonaut is orbiting earth in a spacecraft at an altitude $h = 630$ km with a speed of 8 km/s. If radius of the earth is 6400 km, the acceleration of the cosmonaut is (1) 9.10 m/s ² (2) 9.80 m/s ² (3) 10.0 m/s ² (4) 9.88 m/s ²									

Gravit	ation													
22.	Two satellites A and B go around a planet P in circular orbits having radius 4R and R respectively. If the speed of satellite A is 3v, then the speed of satellite B will be													
	(1) $6V$ (2) $9V$ (3) $3V$					(4) none of these								
Compr	ehens i A pair mass to the	i on # 1 of stars m such size of e	rotates a that M = either sta	about a c 2m. Th r).	commor e distan	n center lice betw	of mass. een the	. One of centres	the stars of the st	has a n ars is d	nass M a (d being	and the c large co	other has compared	
23.ൔ	The period of rotation of the stars about their common (1) $\sqrt{\frac{4\pi^2}{\text{Gm}}\text{d}^3}$ (2) $\sqrt{\frac{8\pi^2}{\text{Gm}}\text{d}^3}$ (3)						ommon (3)	nmon centre of mass (in terms of d, m, G.) is (3) $\sqrt{\frac{2\pi^2}{3Gm}d^3}$ (4) $\sqrt{\frac{4\pi^2}{3Gm}d^3}$						
24.🖎	The ra (1) 1	atio of th	e angula	r mome (2) 2	ntum of	the two	stars ab (3) 4	out their	. commo	n centre (4) 9	of mass	s (L _m / L _N	ı) is	
25.⊾	The ra (1)1	atio of ki	netic ene	rgies of (2) 2	the two	stars (ł	К _т /К _М .) і (3) 4	S		(4) 9				
Compr	ehensi An art of esc at the	i on # 2 ificial sa ape velo surface	tellite is r ocity fron of earth.	noving i n the su (R = 64	n a circu rface of 100 km)	ılar orbit earth. R :	around is the ra	the earth adius of	n with a s earth an	peed ec d g is a	ual to ha	alf the m on due t	agnitude o gravity	
26.	Then 1 (1) 32	the dista 00 km	ince of s	atellite fi (2) 64	rom the 00 km	surface	of earth (3) 12	is 2800 km		(4) 48	(4) 4800 km			
27.	The ti	me perio	od of revo	olution o	f satellit	e in the	given or	bit_is_						
	(1) 2π	$\sqrt{\frac{2R}{g}}$		(2) 2π	$\sqrt{\frac{4R}{g}}$		(3) 2τ	(3) $2\pi \sqrt{\frac{8R}{g}}$ (4) $2\pi \sqrt{\frac{6R}{g}}$						
28.	If the which	satellite it hits th	e is stopp e surfac	ed sude e of the	denly in earth.	its orbit	and all	owed to	fall free	y onto t	he earth	i, the sp	eed with	
,	(1) \sqrt{gR} (2) $\sqrt{1.5gR}$ (3) $\sqrt{\frac{gR}{2}}$ (4) $\sqrt{\frac{gR}{2}}$								gR √2					
	AP	SP /	Ansv	ver	s⊨									
(Þ۸	RT - I							
1.	(4)	2.	(1)	3.	(2)	4.	(3)	5.	(2)	6.	(4)	7.	(2)	
8.	(1)	9.	(2)	10.	(4)	11.	(1)	12.	(2)	13.	(3)	14	(1)	
15.	(2)	16.	(3)	17.	(3)	18.	(1)	19.	(2)	20.	(1)	21.	(3)	
22.	(4)	23.	(3)	24.	(1)	25.	(3)	26.	(1)	27.	(3)	28.	(2)	
29.	(1)	30.	(3)											
	(4)	•		•		PA	RT - II	-		•		-		
1.	(1)	2.	(2)	3.	(2)	4.	(2)	5.	(4)	6. 40	(2)	1.	(4)	
ŏ.	(3)	9.	(3)	10.	(2)	11.	(3)	12.	(3)	13.	(3)	14.	(1)	

16.

23.

(4)

(1)

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(1)

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26.

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(3)

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21.

28.

(1)

(1)